

The behaviour of single
scale hard small x
processes in QCD near
the black disk limit

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(based on

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small x physics meeting

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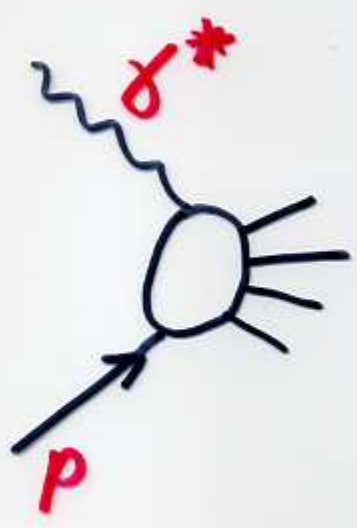
1. Introduction: theoretical challenges of small x physics and instabilities of pQCD approximation.

2. Effective field theory from pQCD, tachyon dynamics and color inflation

3. Color networks

4. Conclusion

I Challenges of small x physics



$$g^2 = -Q^2$$

$$v = 2pQ$$

$$x = \frac{Q^2}{2pQ}$$

moderate x:

$$\alpha_s \ll 1, \quad \frac{\alpha_s}{\pi} \ln \frac{Q^2}{\Lambda_{QCD}^2} \sim 1$$

$$F_2(x, Q^2) = \sum f_i(x, \alpha_s \ln \frac{Q^2}{\Lambda^2}) \alpha_s^i$$

for each f_i - QCD evolution equation.

For small x - new parameter

$$\ln \frac{x_0}{x} - \text{large} \Rightarrow \text{DGLAP}$$

?

4.
One of the striking
predictions of DGLAP
(LO+NLO)

$$F_{2p}, xG_p \sim x^{-\gamma(Q^2)}$$

$\gamma \sim 0.2$ for $Q^2 \sim 10 \text{ GeV}^2$
(consistent with HERA data)

Another approach - BFKL ⁵

$$\alpha_s \ll 1, \quad \alpha_s \ln \frac{x_0}{x} \sim 1$$

$$F_2, xG \sim x^{-\gamma}$$

$$\gamma = 4 \log 2 \frac{\alpha_s (10^2) N_c}{\pi} (1 - \dots)$$

(G. Salam)

However, LO+NLO BFKL
and LO+NLO DGLAP
give the same results
for $x < 10^{-7} \div 10^{-8}$

(G. Salam, M. Ciafaloni, P. Colferai,
A. Stasto 1999, 2002)

Thus in the kinematics up to
small x N²LO calculations

However: problem with unitarity.

6.

$$\sigma_{DGLAP}^{\gamma^* T} (\text{inclusive}) \sim \frac{1}{x^{1/2}(Q^2)} \frac{1}{Q^2}$$

violates Froissart bound
or, in impact parameter space
representation

$$A(\text{dipole} + T \rightarrow \text{dipole} + T) = \\ = 2s \int d^2b e^{i\vec{q}_t \cdot \vec{b}} f(b, s, Q^2)$$

$$|f(b, s, Q^2)| \leq 1 \quad \underline{\text{must be,}}$$

but reaches

$$f \sim 1 \quad \text{for } Q^2 \sim 10 \text{ GeV}^2$$

already at $x \sim 10^{-4} - 10^{-5}$

(LHC kinematics)

$$\sigma_{tot} = 2\pi b_{max}^2$$

$$b_{max} \sim \frac{1}{\gamma} \ln s/s_0$$

(Froissart Bound)

How to achieve such behaviour in perturbative QCD?

The observation —

- α_s series are only asymptotical and defined up to $O(\exp(-\frac{1}{2s}))$ terms for sufficiently large energies the latter terms

dominate, α_s expansion nonvalid — black disk limit (Froissart)

behaviour?

